

1 **The Spatio-temporal Variation Pattern of**
2 **Thermospheric Mass Density Revealed by**
3 **Co-clustering**

4 **Zhaoyuan Yu^{1,2,3}, Zhenxia Liu¹, Zengjie Wang¹, Lingling Sun¹, Hong Gao⁴,**
5 **Wen Luo^{1,2,3}, Linwang Yuan^{*1,2,3},**

6 ¹Ministry of Education, Key Laboratory of Virtual Geographic Environment (Nanjing Normal
7 University), Nanjing, China;

8 ²Jiangsu Center for Collaborative Innovation in Geographical Information Resource Development and
9 Application, Nanjing, China;

10 ³State Key Laboratory Cultivation Base of Geographical Environment Evolution (Jiangsu Province),
11 Nanjing, China;

12 ⁴Faculty of Geomatics, Lanzhou Jiaotong University, Lanzhou, Gansu Province, China;

13 **Key Points:**

- 14 • Analysis of the spatio-temporal structural characteristics of thermospheric mass
15 density variations using GOCE satellite data
16 • Exploration on the structures of spatial changes of thermospheric mass density
17 in long time series using a co-clustering method
18 • Comparative analysis of the influence between different temporal and spatial struc-
19 tural characteristics of thermospheric mass density

Corresponding author: Linwang Yuan, yuanlinwang@njnu.edu.cn

Abstract

The spatio-temporal distribution characteristics of thermospheric mass density have been given more attention with an increasing demand for spacecraft launches and low Earth orbital prediction. More and more patterns of spatial structure and temporal variation are being discovered. Notwithstanding these developments, the study of spatio-temporal coupling in characteristics analysis remains quite limited. In this study, we use a co-clustering method to explore and analyze the spatio-temporal coupling structural characteristics of thermospheric mass density. The processed GOCE satellite dataset is divided into 5 temporal clusters and 20 spatial clusters by the co-clustering method. In terms of spatial structure, the density has an obvious zonal distribution structure and hemispheric asymmetry. Moreover, due to the influence of the Earth's magnetic field, there is an average angle about 2.00° between the band structure and the latitudinal circle. In terms of temporal structure, the temporal patterns of density can be grouped into five period types, namely the quiet period, the moderate activity period, the event period, the oscillation period and the recovery period. And significant positive correlation can be found between the F10.7 indices and the temporal density variation. This study explores the spatial structure and temporal pattern of thermospheric mass density and its driving forces from the perspective of spatio-temporal coupling based on a statistical method, which can provide a new idea of spatio-temporal coupling method for spatio-temporal evolution of thermospheric mass density.

1 Introduction

Thermospheric mass density has a significant impact on spacecraft launches and low Earth orbital prediction since it is the primary source of the atmospheric drag (Emmert, 2015; Doornbos et al., 2008), which has complex spatio-temporal variations and response characteristics commonly driven by solar ultraviolet irradiance, solar wind, energetic particles from the magnetosphere, and waves originating from the lower atmosphere (P. Wang et al., 2022). It is essential to better analyze and explore the spatio-temporal characteristics of the thermospheric mass density to guide the space exploration.

With continuous research on the observational data of the thermospheric, the variation pattern of thermospheric mass density is not only related to temporal changes (The seasonal changes, the semiannual changes, daily changes, etc.) but also closely related to spatial changes (Changes in latitude and longitude, changes in polar regions), as well as anomalies such as thermospheric storms, night enhancement, equatorial anomalies, longitude distribution, and irregular polar structures (Emmert, 2015; Liu et al., 2007; Qian & Solomon, 2012). In terms of spatial structure, Walterscheid et al. (Walterscheid & Crowley, 2015) studied the spatial structure of thermospheric mass density during geomagnetic calm and active periods using the TMEGCM model and provided a detailed interpretation of its low and high-density regions. Thermospheric mass density exhibits hemispheric asymmetry (Meier et al., 2015), and overall thermospheric mass density exhibits a slightly higher structure in the southern hemisphere than in the northern hemisphere, and the maximum value occurs in the equatorial region. In terms of temporal pattern, related studies (Lei et al., 2011; Danilov et al., 1992; Haris et al., 1994; Weng et al., 2018) show the significant variation in atmospheric mass density with diurnal, seasonal and annual/semi-annual variation, especially, both amplitude and phase of the thermospheric mass density seasonal variations have strong solar activity dependences (Weng et al., 2018). Many scholars have used empirical models and spatial statistics to reveal the characteristics of the temporal variation of the thermospheric (Yuan et al., 2019; Mehta et al., 2017; Ren et al., 2021), but the temporal and spatial characteristics are often not closely linked in these analyses. Considering the universal connection between temporal and spatial features within geographic phenomena (Deng et al., 2013), the density distribution characteristics of spatial and temporal are not independent but coupled with each other. Therefore, it is especially important to adopt an analytical method that can

72 establish the coupling of temporal and spatial features to analyze the temporal and spa-
 73 tial evolution characteristics of thermospheric atmospheric density.

74 As one of the common methods to study the spatial and temporal differentiation
 75 of geographical phenomena, clustering analysis groups the data with similarities to iden-
 76 tify the internal structure and numerical divergence patterns of data sets, and discov-
 77 ers the divergence pattern within the data(Andrienko et al., 2009), it is particularly ef-
 78 fective for exploring the hidden features within geographical phenomena. Clustering meth-
 79 ods are particularly useful in atmospheric and space weather studies because the results
 80 of clustering over large regions and long time scales are more practical than those sum-
 81 marized from short time observations of several locations(Zirlewagen & Wilpert, 2010).
 82 The spatial clustering and temporal clustering methods have been used by many schol-
 83 ars to study the atmospheric spatial and temporal divergence patterns. For example, Wang
 84 et al. investigated the spatial and temporal patterns of atmospheric ozone pollution by
 85 Spatio-temporal Shared Nearest Neighbor clustering algorithm(S. Wang et al., 2013), Zhao
 86 et al.used clustering to explore anomaly correlation and to reveal how global climate model(GCM)
 87 forecasts perform as time progresses(Zhao et al., 2017). Ma et al. clustered potential flare-
 88 ing active regions by applying Distance Density clustering on individual parameters and
 89 further organized the clustering results into a multivariate time series decision tree and
 90 addressed the problem of flare prediction(Ma et al., 2017). However, the spatial diver-
 91 gence patterns obtained using spatial clustering methods alone do not reveal time-varying
 92 behavior, and vice versa. Thus, a clustering method that can obtain spatial and tem-
 93 poral divergence patterns is needed to explore the spatial and temporal coupling char-
 94 acteristics. Co-clustering is able to show its variation with time and space by perform-
 95 ing both temporal and spatial clustering of the thermospheric atmospheric density, es-
 96 tablishing its coupling characteristics in both temporal and spatial dimensions. There-
 97 fore, we use a co-clustering algorithm to cluster thermospheric mass density data, and
 98 study the spatio-temporal structure and patterns of thermospheric mass density under
 99 the interaction and interplay of time and space.

100 In the first section, we systematically analyze changes in spatio-temporal charac-
 101 teristics of thermospheric mass density. The second section briefly introduces the data
 102 sources and overall framework of the methods. The third section analyzes the tempo-
 103 ral and spatial characteristics of thermospheric mass density based on the spatio-temporal
 104 co-clustering, and discusses the dynamic evolution of spatial structure and temporal pat-
 105 terns and driving force based on the co-cluster result. Finally, we summarized conclu-
 106 sion, and explained the feasibility of our results.

107 2 Data and Method

108 2.1 GOCE Data

109 GOCE (Gravity field and steady-state ocean circulation explore) satellite thermo-
 110 spheric mass density data with better observation continuity and quality, have been used
 111 to study the characteristics of gravity wave distribution at various scales and the sea-
 112 sonal variation distribution pattern, increasing our understanding of the variation pat-
 113 tern of the thermospheric mass density (Forbes et al., 2016; Liu et al., 2017). We use the
 114 GOCE satellite data from November 1, 2009, to October 20, 2013, including 1450 days.
 115 Its latitude coverage is from 83. 5° S to 83° N and the average altitude naturalized to
 116 270km altitude. Due to the inconsistency between the sampling interval of the satellite
 117 and the revisit period of the satellite, it is difficult to have multiple consecutive data at
 118 the same location, which results in extremely sparse satellite data in space. So we di-
 119 vide the globe into a 2°x2° grid so as to organize discrete GOCE satellite data regularly,
 120 and treat one day of sample points within the same grid as one data set. If there are mul-
 121 tiple sample points on the same grid, the average value of the sampling points is taken
 122 as thermospheric mass density. In addition, since the thermosphere has a thickness, the
 123 whole thermosphere is a three-dimensional space. We stretch the three-dimensional ther-

124 mospheric space into a two-dimensional thermospheric plane to study the horizontal struc-
 125 tural characteristics of the thermospheric mass density. Then, we divide the globe into
 126 16,200 grids by unfolding the two-dimensional plane as a $2^\circ \times 2^\circ$ grid to achieve one-dimensional
 127 spatial properties. Finally, one day as a basic unit in temporal dimension. In this way,
 128 the dimensional division of spatio-temporal co-clustering is achieved.

129 **2.2 Overall framework of the methods**

130 After preprocessing the data, the data are analyzed using co-clustering to explore
 131 the internal spatial and temporal characteristics, and divided the spatial and temporal
 132 clusters. Then, geodetector is used to determine the most appropriate number of clus-
 133 ters. For the co-clustering results, they are spread spatially to explore their spatial char-
 134 acteristics and the angle of the zonal distribution structure, and spread temporally to
 135 explore the temporal characteristics and compared with the relevant space weather in-
 136 dices.

137 **2.2.1 Co-clustering**

138 Co-clustering is used to explore the internal spatial and temporal characteristics,
 139 and divided the spatial and temporal clusters. Clustering analysis is an important un-
 140 supervised learning method. The essential idea of clustering analysis is to divide data
 141 objects with similarities into one group. It divides the data set into several groups by
 142 a certain criterion (such as distance or similarity of objects), and each group is called
 143 a cluster. Traditional clustering methods can be divided into two types: one is to clus-
 144 ter samples based on all variables, and the other is to cluster variables based on all samples(Gore,
 145 2000). These two types of methods only cluster samples or variables, so they are often
 146 referred to as unidirectional or one-way clustering. Different from one-way clustering,co-
 147 clustering is the clustering of data blocks (subsets of data), which is based on the sim-
 148 ilarity of element values within the data matrix, forming a sub-matrix partitioning scheme
 149 so that the sub-matrices elements are as similar as possible, thus achieving simultane-
 150 ous clustering in both spatial and temporal dimensions.

151 Both temporal clustering and spatial clustering are one-way clustering, which is only
 152 based on a certain dimension. Figure 1(a) shows thermospheric mass density example
 153 data (The values do not represent the real density), Figure 1(b) treats space as an ob-
 154 ject and time as its property to obtain a spatial group with similarity in time, Figure
 155 1(c) treats time as an object and space as its property to obtain temporal group with
 156 similarity in space. Unlike one-way clustering, two-way clustering can cluster both time
 157 and space, dividing data that are similar in both time and space into the same spatio-
 158 temporal group, in which both spatial and temporal values are similar, as in Figure 1(d).
 159 Due to the sparseness of data, the clustering method we used is Biclustermd (Li et al.,
 160 2020; Reisner et al., 2019), which is suitable for datasets with missing values and proved
 161 to have good performance for very sparse data in an agricultural application and a movie
 162 rater application(Li et al., 2020).

163 **2.2.2 Geodetector**

164 Geodetector is used to determine the most appropriate number of clusters. Geode-
 165 tector is a statistical method to detect spatial dissimilarity and to reveal the driving forces
 166 behind it(J. Wang & Xu, 2017), which has been widely used in natural sciences, social
 167 sciences, environmental sciences and human health(Dong et al., 2017; Ju et al., 2016; Gao
 168 et al., 2016). In this study, the reliability of co-clustering results is assessed by the q -statistic

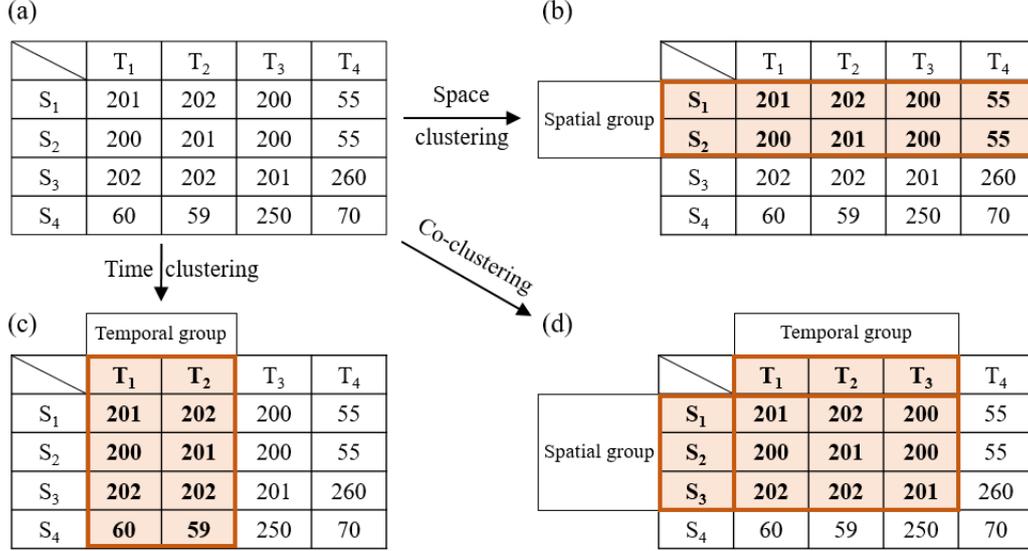


Figure 1. Schematic diagram of co-clustering (Modified from (Wu et al., 2020)). (a) Thermospheric mass density data matrix, T_i denotes the time dimension, S_i denotes the spatial dimension. (b) A group with temporal similarity obtained by spatial clustering. (c) A group with spatial similarity obtained by temporal clustering. (d) Co-clustering to obtain spatio-temporal groups with similarity in both time and space dimension.

169 in geodetector to measure, and the expression is as follow:

$$q = 1 - \frac{\sum_{h=1}^L N_h \sigma_h^2}{N \sigma^2}$$

170 Where: L is the strata of variables, that is cluster in this paper; N_h and N are the
 171 number of units in strata h and the whole area, respectively; σ_h^2 and σ^2 are the squares
 172 in strata h and the whole area, respectively. The value range of q is $[0,1]$. Larger value
 173 of q indicates more pronounced heterogeneity.

174 2.2.3 Calculation of angle of zonal distribution structure

175 The process of calculating the angle of zonal distribution structure is divided into
 176 four main steps. Firstly, for structures composed of clusters of the same spatial clusters,
 177 the average of the latitudes on the same longitude is obtained to get the center point $A(X_{lon}, Y_{lat})$.
 178 Secondly, binomial fitting is carried out with the center points of these different longi-
 179 tudes. The vertex of binomial is the lowest point or the highest point of the whole struc-
 180 ture, and the plane passing through this vertex is the axial plane. Thirdly, these centers
 181 are projected on the plane where the axis plane is located, and their corresponding
 182 positions on the plane are obtained. For a given central point $A(X_{lon}, Y_{lat})$, the equa-
 183 tion for its projection onto the axis plane is as follows:

$$\begin{cases} X = R * (1 - \cos X_{lon}) \\ Y = R * \sin Y_{lat} \end{cases}$$

184 Where X_{lon} is the longitude of the center point A; Y_{lat} is the latitude of the center point
 185 A; X is the horizontal coordinate of the projection of the center point A on the axis plane;

186 Y is the vertical coordinate of the projection of the center point A on the axis plane; X_0
 187 is the longitude of the longitude circle where the axis plane is located; R is the radius
 188 of the earth. Finally, the slope of the projection of the zonal distribution structure on
 189 the axial plane is fitted according to the projected coordinates, and the angle is calcu-
 190 lated by the slope. Considering that the zonal distribution structure is symmetrical about
 191 the axis plane, each zonal distribution structure needs to find the inclination angle on
 192 both sides of the axis plane by this method, and the average value of the inclination an-
 193 gle on both sides is used as the inclination angle of this zonal distribution structure.

194 3 Results and Discussion

195 3.1 Spatio-temporal pattern

196 Based on the pre-processed two-dimensional spatio-temporal dataset of thermospheric
 197 mass density, the data is organized into a data matrix, where the rows represent total
 198 14940 (83×180) grids for the spatial dimension and the columns represent 1450 days
 199 for the temporal dimension. The analysis is then performed using a co-clustering method
 200 in Section 2.2. To determine the spatial clusters and temporal clusters, clustering anal-
 201 ysis is performed using the number of spatial clusters and temporal clusters with step
 202 lengths of 2 from 10 to 40 and 1 from 3 to 7, respectively. The clustering results are eval-
 203 uated using the geodetector q -value in Section 2.3. Finally, the results show that the clus-
 204 tering effect is better when the temporal clusters is 5 and the spatial clusters is 20. At
 205 this time, the q -value of the geodetector is 0.7948, indicating that the result has good
 206 differentiation(Dong et al., 2017).

207 The co-clustering results are shown in Figure 2(a). The horizontal axis represents
 208 14940 grids in the dataset, arranged in the order of belonging to the same spatial clus-
 209 ters, the vertical axis represents the 1450 days in the dataset, arranged in the order of
 210 belonging to the same temporal clusters. The black solid lines represent the boundary
 211 of different spatial clusters and temporal clusters. And the color indicates the mass den-
 212 sity value, especially, white means the data is empty because of data sparsity. It can be
 213 seen that the density values of the same cluster are relatively close, the boundaries be-
 214 tween different clusters are relatively obvious, which means that co-clustering analysis
 215 can be a good way to construct spatial and temporal characteristics of thermospheric
 216 mass density considering the sparsity of the data. Since both temporal and spatial pa-
 217 rameters are considered in the clustering calculation process, the results of co-clustering
 218 well establish the coupling relationship between temporal and spatial dimensions, which
 219 is obviously different from the traditional study of thermospheric mass density in which
 220 only its spatial structure or temporal characteristics are considered alone. It solves the
 221 shortcoming of uncoupled spatio-temporal characteristics in the traditional analysis, which
 222 is significant for the construction of a spatio-temporal coupled thermospheric analysis
 223 model and discovering more structural characteristics about the thermospheric mass den-
 224 sity from the perspective of spatio-temporal coupling.

225 The results of spatio-temporal co-clustering in Figure 2(a) are expanded in space
 226 to obtain the spatial pattern and evolutionary dynamics of thermospheric mass density
 227 as shown in Figure 2(b). The spatial structure of thermospheric mass density at differ-
 228 ent time periods in Figure 2(b) all show obvious latitudinal divergence, that is, the spa-
 229 tial structure is dominated by latitudinal variation and shows the overall characteristics
 230 of higher density at the poles and lower density at the equator. And the density change
 231 in the southern hemisphere is more pronounced than that in the northern hemisphere.

232 The results of spatio-temporal co-clustering in Figure 2(a) are expanded in time,
 233 and the temporal clustering change curves are plotted according to the five temporal clus-
 234 ters, as shown in Figure 2(c). Curve A represents the original temporal clustering results
 235 in co-clustering, and curve B is obtained by removing outliers which is larger differences
 236 in cluster with adjacent time, and the same height indicates the same groups, which the
 237 height from low to high is Cluster1, Cluster2, Cluster3, Cluster4, Cluster5. Since the pur-

pose of co-clustering is to measure the similarity of the density in the temporal dimension and the spatial dimension at the same time, so the group of time does not show continuous segmentation under the influence of space. Figure 2(c) shows the short-period variation in thermospheric mass density over time periods.

Throughout the study period, there are some series of time periods with overall high thermospheric mass density, which indicates the existence of external driving forces to make the density change with time. And from the spatial perspective, there are obvious spatially divergent characteristics of the thermospheric atmospheric density. In the following, we will reveal its structural features from the spatial and temporal dimensions based on the results of co-clustering.

3.2 Spatial structure and dynamic evolution

The co-clustering method aggregates the density values with spatial-temporal similarity into the same group. The result of Figure 2(b) well reflects the macroscopic spatial structure characteristics of the thermospheric atmosphere at long time scales, i.e., thermospheric mass density shows hemispheric asymmetry, thermospheric mass density in the southern hemisphere is slightly higher than that in the northern hemisphere, and the maximum value occurs in the equatorial region (Meier et al., 2015). This conclusion is similar to many studies on the spatio-temporal pattern of the thermospheric atmosphere, such as zonal distribution structure (Emmert, 2015; Liang et al., 2015), hemispheric asymmetry (Meier et al., 2015), and so on.

It is obvious through Figure 2(b) that there is an angle between the zonal distribution structure of thermospheric mass density and latitudinal circle. Based on this idea, the angle between the approximate axis of the zonal distribution structure and latitudinal circle is obtained by the method in section 2.4, and the results are shown in Table 1 (considering the deformation in polar regions, only the complete zonal distribution structure in the range of 60°N to 60°S is selected here). The results of Table 1 show that the multiple zonal distribution structure all exhibit a certain angle of entrainment (average angle about 2.00°), and the axial plane tends to fluctuate around the line from 12°W to 168°E. This phenomenon is close to the dip of the Earth's magnetic axis (70°W to 110°E.), it is statistically significant to show that the spatial density distribution of the thermospheric is influenced not only by various types of solar activity, but also by the Earth's own magnetic field. This result is coupled with studies on the Earth's magnetic field (Laundal et al., 2017; Lei et al., 2012). Not only can the thermospheric influence the distribution of the Earth's magnetic field through solar wind and ionization, but the Earth's magnetic field can also influence the density characteristics of thermospheric by acting on charged particles.

3.3 Temporal patterns and driving force

Based on Figure 2(c), it can be concluded that thermospheric mass density shows a periodic variation under the action of external driving forces. It has been shown by related studies that thermospheric mass density is closely related to changes in factors such as solar activity and geomagnetic activity, such as solar EUV and UV radiation, high-energy particle deposition in the polar regions, magnetospheric plasma convection, and various atmospheres uploaded from the middle and lower atmospheres to the upper atmosphere (Rostoker et al., 1998; Baker & Kanekal, 2008; Liu & Lühr, 2005). This external energy will heat the thermospheric atmosphere, causing the ionization and decomposition of various components in the thermospheric, and then driving the changes in thermospheric mass density.

F10.7 and Ap are two typical space weather indices that are often considered as important driving force of thermospheric changes, so it is important to explore the relationship between these two indices and the co-clustering results. The joint analysis of the density clustering characteristic curves and the associated space weather indices resulted in Figure 3 and Table 2, Table 3. From the Figure 3, it can be seen that the tem-

Table 1. The angle between the approximate axis of the zonal distribution structure and the latitudinal circle

Cluster	Angle (°)	Axial Plane
<i>Cluster2N</i>	2.74	29° W to 151° E
<i>Cluster2S</i>	2.01	9° W to 171° E
<i>Cluster3S</i>	1.95	7° W to 173° E
<i>Cluster4N</i>	1.81	3° W to 177° E
<i>Cluster4S</i>	1.43	5° W to 175° E
<i>Cluster5S</i>	1.62	7° W to 173° E
<i>Cluster7S</i>	1.31	15° W to 165° E
<i>Cluster8N</i>	3.94	27° W to 153° E
<i>Cluster9S</i>	1.15	3° W to 177° E
<i>Average</i>	2.00	12° W to 168° E

290 poral clustering curve and the F10.7 curve fit well, when the density is at Cluster1, F10.7
 291 is low overall, at Cluster2, F10.7 is moderate overall, when Cluster3, F10.7 is high over-
 292 all, when Cluster4, F10.7 is in a fluctuating period overall and fluctuate in a wide range,
 293 while Cluster5 is in a time series that belongs to a period immediately following the large
 294 fluctuation of Cluster4 after the small fluctuation period. Taking the clustering results
 295 of the three classes Cluster1, Cluster2, and Cluster3 (Considering that the values in clus-
 296 ter4 and cluster5 fluctuate widely and represent a fluctuating period, only the cluster1,
 297 cluster2, cluster3 that represent different densities are considered when establishing the
 298 correlation link) characterizing the density size and the F10.7 curve for correlation anal-
 299 ysis, the Pearson correlation coefficient is calculated to obtain $p = 0.755$, it suggests a
 300 significant positive correlation between the temporal patterns of thermospheric mass den-
 301 sity and F10.7. The curve characteristics shown in Figure 3 are also more consistent with
 302 the statistical results of F10.7 in Table 2. This series of features indicates that the tem-
 303 poral patterns of thermospheric mass density are strongly influenced by the solar activ-
 304 ity on the time scale, and the overall temporal density characteristics are closely related
 305 to the variation of the F10.7, and to some extent, the variation pattern of the F10.7 in-
 306 dex can be used to characterize the temporal density variation pattern of the thermo-
 307 spheric mass density.

308 According to the comparing results and the intensity of the solar activity repre-
 309 sented by F10.7, the five temporal clusters cluster1, cluster2, cluster3, cluster4, and clus-
 310 ter5 of the thermospheric mass density can be classified into the quiet period, the mod-
 311 erately active period, the event period, the oscillation period and the recovery period,
 312 respectively, which can characterize well the effect of solar activity on thermospheric mass
 313 density. This shows that co-clustering can well analyze the temporal characteristics of
 314 thermospheric mass density.

315 Combining Figure 3 and Table 3, it can be found that there is no obvious correla-
 316 tion between the clustering result and Ap, where $p = 0.065$ (Taking the clustering re-
 317 sults of the three classes Cluster1, Cluster2, and Cluster3 characterizing the density size
 318 and the AP curve for correlation analysis). This shows that geomagnetic activity has no
 319 direct influence on the temporal variation of the thermospheric mass density. Combined
 320 with spatial analysis, it can be seen that geomagnetic field has obvious influence on the
 321 spatial pattern of characteristics in the thermospheric mass density, but it has no ob-
 322 vious driving effect on the density change in temporal scale.

Table 2. F10.7 intensive statistics under different clusters

Cluster	Total (d)	low	Moderate	High	Very High
<i>C1</i>	577	252	280	45	0
<i>C2</i>	366	0	55	290	21
<i>C3</i>	273	0	27	212	34
<i>C4</i>	92	0	7	82	3
<i>C5</i>	142	0	17	125	0

Table 3. AP statistics under different clusters

Cluster	Total (d)	Quiet	Unsettled	Active	Minor Storm	Major Storm
<i>C1</i>	577	442	101	28	5	1
<i>C2</i>	366	205	89	55	15	2
<i>C3</i>	273	211	40	14	8	0
<i>C4</i>	92	61	18	10	2	1
<i>C5</i>	142	84	32	23	0	3

4 Conclusion

In this paper, the spatio-temporal coupling characteristics of thermospheric mass density are analyzed from the spatio-temporal correlation of natural phenomena features, using co-clustering method considering both from the temporal and spatial dimensions, exploring dynamic evolution of spatial structure and driving force of temporal patterns. Temporal structural characteristics are closely related to solar activity. And temporal patterns of thermospheric mass density can be concluded into five periods, namely the quiet period, the moderate activity period, the event period, the oscillation period and the recovery period, which can characterize well the effect of solar activity on thermospheric mass density. Especially, there is a significant positive correlation between the F10.7 indices and the density temporal variation. And the F10.7 indices can be used to some extent to characterize the global thermospheric mass density, considering the difficulty of thermospheric mass density measurement. The spatial structural characteristics show obvious zonal distribution structure, in which the equatorial region and the middle and low latitudes of northern latitude have obvious structures and boundaries, but there is no obvious structure in the southern hemisphere and the poles. There is an average angle about 2.00° between the band structure and the latitudinal circle because of the influence of the Earth's magnetic field. Not only can the thermospheric influence the distribution of the Earth's magnetic field through solar wind and ionization, but the Earth's magnetic field can also influence the density characteristics of thermospheric by acting on charged particles.

At present, this research is only a rough exploration and analysis of the spatio-temporal structure of thermospheric mass density on a large scale. Since the statistical method from data is used in this paper, it is not possible to explain the inner formation mechanism of statistical phenomena, such as the dip angle phenomenon of the band structure. There are still many factors that have not been considered, such as the influence of different parameters (geomagnetic, gravity waves). This paper only analyzes the temporal and spatial characteristics of thermospheric mass density from a holistic perspective, and some local anomalous details need to be further explored. Many spatio-temporal

352 phenomena and spatio-temporal structures are still on the surface, and their internal mech-
 353 anism needs further explanation.

354 5 Data Availability Statement

355 Thermospheric mass density data of GOCE used in the present paper are from TUDelft
 356 (<ftp://thermosphere.tudelft.nl/>)

357 Acknowledgments

358 ZYY and ZXL contributed equally to this work and should be considered co-first authors.
 359 This work was supported in part by the National Natural Science Foundation of China
 360 under Grant 42130103, 41976186, 41971404, and the Postgraduate Research & Practice
 361 Innovation Program of Jiangsu Province under Grant KYCX22_1578.

362 References

- 363 Andrienko, G. L., Andrienko, N. V., Rinzivillo, S., Nanni, M., & Giannotti, F.
 364 (2009). Interactive visual clustering of large collections of trajectories. In
 365 *2009 IEEE Symposium on Visual Analytics Science and Technology* (pp. 3–10).
 366 doi: <https://doi.org/10.1109/VAST.2009.5332584>
- 367 Baker, D., & Kanekal, S. (2008). Solar cycle changes, geomagnetic varia-
 368 tions, and energetic particle properties in the inner magnetosphere. *Jour-
 369 nal of Atmospheric and Solar-Terrestrial Physics*, *70*(2-4), 195–206. doi:
 370 <https://doi.org/10.1016/j.jastp.2007.08.031>
- 371 Danilov, A., Kalgin, Y. A., & Pokhunkov, A. (1992). Seasonal and latitudinal vari-
 372 ations of turbulence in the lower thermosphere. *Advances in Space Research*,
 373 *12*(10), 123–129. doi: [https://doi.org/10.1016/0273-1177\(92\)90453-5](https://doi.org/10.1016/0273-1177(92)90453-5)
- 374 Deng, M., Liu, Q., Wang, J., & Shi, Y. (2013). A general method of spatio-temporal
 375 clustering analysis. *Science China Information Sciences*, *56*(10), 1–14.
- 376 Dong, Y., Qian, X., Qian, X., Chengdong, X., & Yuying, W. (2017). Delineation of
 377 the northern border of the tropical zone of china’s mainland using geodetector.
 378 *Acta Geographica Sinica*, *72*(1), 135.
- 379 Doornbos, E., Klinkrad, H., & Visser, P. (2008). Use of two-line element data for
 380 thermosphere neutral density model calibration. *Advances in Space Research*,
 381 *41*(7), 1115–1122.
- 382 Emmert, J. T. (2015). Thermospheric mass density: A review. *Advances in Space
 383 Research*, *56*(5), 773–824. doi: <https://doi.org/10.1016/j.asr.2015.05.038>
- 384 Forbes, J. M., Bruinsma, S. L., Doornbos, E., & Zhang, X. (2016). Gravity wave-
 385 induced variability of the middle thermosphere. *Journal of Geophysical Re-
 386 search: Space Physics*, *121*(7), 6914–6923. doi: [https://doi.org/10.1002/
 387 2016JA022923](https://doi.org/10.1002/2016JA022923)
- 388 Gao, F., Li, S., Tan, Z., Wu, Z., Zhang, X., Huang, G., & Huang, Z. (2016). Un-
 389 derstanding the modifiable areal unit problem in dockless bike sharing usage
 390 and exploring the interactive effects of built environment factors. *Interna-
 391 tional Journal of Geographical Information Science*, *35*(9), 1905–1925. doi:
 392 <https://doi.org/10.1080/13658816.2020.1863410>
- 393 Gore, P. A. (2000). 11 - cluster analysis. In H. E. Tinsley & S. D. Brown
 394 (Eds.), *Handbook of applied multivariate statistics and mathematical mod-
 395 eling* (p. 297–321). San Diego: Academic Press. Retrieved from [https://
 396 www.sciencedirect.com/science/article/pii/B9780126913606500124](https://www.sciencedirect.com/science/article/pii/B9780126913606500124)
 397 doi: <https://doi.org/10.1016/B978-012691360-6/50012-4>
- 398 Haris, P. A. T., Stevens, T. D., Maruvada, S., & Philbrick, C. R. (1994). Latitudinal
 399 variation of middle atmospheric density and temperature. *Advances in Space
 400 Research*, *14*(9), 83–87. doi: [https://doi.org/10.1016/0273-1177\(94\)90120-1](https://doi.org/10.1016/0273-1177(94)90120-1)

- 401 Ju, H., Zhang, Z., Zuo, L., Wang, J., Zhang, S., Wang, X., & Zhao, X. (2016).
 402 Driving forces and their interactions of built-up land expansion based on
 403 the geographical detector – a case study of Beijing, China. *International*
 404 *Journal of Geographical Information Science*, *30*(11), 2188–2207. doi:
 405 <https://doi.org/10.1080/13658816.2016.1165228>
- 406 Laundal, K. M., Cnossen, I., Milan, S. E., Haaland, S., Coxon, J., Pedatella, N.,
 407 ... Reistad, J. P. (2017). North–south asymmetries in Earth’s magnetic
 408 field. *Space Science Reviews*, *206*(1), 225–257. doi: <https://doi.org/10.1007/s11214-016-0273-0>
- 410 Lei, J., Forbes, J. M., Liu, H.-L., Dou, X., Xue, X., Li, T., & Luan, X. (2011). Lat-
 411 tudinal variations of middle thermosphere: Observations and modeling. *Jour-
 412 nal of Geophysical Research: Space Physics*, *116*(A12), A12306. doi: <https://doi.org/10.1029/2011JA017067>
- 414 Lei, J., Thayer, J. P., Wang, W., Richmond, A. D., Roble, R., Luan, X., ... Li, T.
 415 (2012). Simulations of the equatorial thermosphere anomaly: Field-aligned ion
 416 drag effect. *Journal of Geophysical Research: Space Physics*, *117*(A1). doi:
 417 <https://doi.org/10.1029/2011JA017114>
- 418 Li, J., Reisner, J., Pham, H., Olafsson, S., & Vardeman, S. (2020). Biclustering with
 419 missing data. *Information Sciences*, *510*, 304–316. doi: <https://doi.org/10.1016/j.ins.2019.09.047>
- 421 Liang, J., Donovan, E., Nishimura, Y., Yang, B., Spanswick, E., Asamura, K., ...
 422 Redmon, R. (2015). Low-energy ion precipitation structures associated with
 423 pulsating auroral patches. *Journal of Geophysical Research: Space Physics*,
 424 *120*(7), 5408–5431. doi: <https://doi.org/10.1002/2015JA021094>
- 425 Liu, H., & Lühr, H. (2005). Strong disturbance of the upper thermospheric den-
 426 sity due to magnetic storms: Champ observations. *Journal of Geophysical Re-
 427 search: Space Physics*, *110*(A9). doi: <https://doi.org/10.1029/2004JA010908>
- 428 Liu, H., Lühr, H., & Watanabe, S. (2007). Climatology of the equatorial thermo-
 429 spheric mass density anomaly. *Journal of Geophysical Research: Space Physics*,
 430 *112*(A5), A05305. doi: <https://doi.org/10.1029/2006JA012199>
- 431 Liu, H., Pedatella, N., & Hocke, K. (2017). Medium-scale gravity wave activity
 432 in the bottomside F region in tropical regions. *Geophysical Research Letters*,
 433 *44*(14), 7099–7105. doi: <https://doi.org/10.1002/2017GL073855>
- 434 Ma, R., Boubrahimi, S. F., Hamdi, S. M., & Angryk, R. A. (2017). Solar
 435 flare prediction using multivariate time series decision trees. In *2017*
 436 *IEEE International Conference on Big Data (Big Data)* (pp. 2569–2578). doi:
 437 <https://doi.org/10.1109/BigData.2017.8258216>
- 438 Mehta, P. M., Walker, A. C., Sutton, E. K., & Godinez, H. C. (2017). New den-
 439 sity estimates derived using accelerometers on board the Champ and Grace
 440 satellites. *Space Weather*, *15*(4), 558–576. doi: <https://doi.org/10.1002/2016SW001562>
- 442 Meier, R. R., Picone, J. M., Drob, D., Bishop, J., Emmert, J. T., Lean, J. L., ...
 443 others (2015). Remote sensing of Earth’s limb by TIMED/GUVI: Retrieval of
 444 thermospheric composition and temperature. *Earth and Space Science*, *2*(1),
 445 1–37. doi: <https://doi.org/10.1002/2014EA000035>
- 446 Qian, L., & Solomon, S. C. (2012). Thermospheric density: An overview of temporal
 447 and spatial variations. *Space Science Reviews*, *168*(1), 147–173. doi: <https://doi.org/10.1007/s11214-011-9810-z>
- 449 Reisner, J., Pham, H., Olafsson, S., Vardeman, S. B., & Li, J. (2019). biclustermmd:
 450 An R package for biclustering with missing values. *R J.*, *11*(2), 69.
- 451 Ren, D., Lei, J., Wang, W., Burns, A., & Luan, X. (2021). Observations and simu-
 452 lations of the peak response time of thermospheric mass density to the 27-day
 453 solar EUV flux variation. *Journal of Geophysical Research: Space Physics*,
 454 *126*(2), e2020JA028756.
- 455 Rostoker, G., Skone, S., & Baker, D. N. (1998). On the origin of relativistic elec-

- 456 trons in the magnetosphere associated with some geomagnetic storms. *Geo-*
457 *physical Research Letters*, 25(19), 3701–3704. doi: [https://doi.org/10.1029/](https://doi.org/10.1029/98GL02801)
458 [98GL02801](https://doi.org/10.1029/98GL02801)
- 459 Walterscheid, R. L., & Crowley, G. (2015). Thermal cell structures in the high-
460 latitude thermosphere induced by ion drag. *Journal of Geophysical Re-*
461 *search: Space Physics*, 120(8), 6837–6850. doi: [https://doi.org/10.1002/](https://doi.org/10.1002/2015JA021122)
462 [2015JA021122](https://doi.org/10.1002/2015JA021122)
- 463 Wang, J., & Xu, C. (2017). Geodetector: Principle and prospective. *Acta Geograph-*
464 *ica Sinica*, 72(1), 19.
- 465 Wang, P., Chen, Z., Deng, X., Wang, J., Tang, R., Li, H., . . . Wu, Z. (2022). The
466 prediction of storm-time thermospheric mass density by lstm-based ensemble
467 learning. *Space Weather*, 20(3), e2021SW002950.
- 468 Wang, S., Cai, T., & Eick, C. F. (2013). New spatiotemporal clustering algo-
469 rithms and their applications to ozone pollution. In *2013 IEEE 13th in-*
470 *ternational conference on data mining workshops* (p. 1061-1068). doi:
471 <https://doi.org/10.1109/ICDMW.2013.14>
- 472 Weng, L., Lei, J., Doornbos, E., Fang, H., & Dou, X. (2018). Seasonal variations
473 of thermospheric mass density at dawn/dusk from goce observations. *Annales*
474 *Geophysicae*, 36(2), 489–496. doi: <https://doi.org/10.5194/angeo-36-489-2018>
- 475 Wu, X., Cheng, C., Qiao, C., & Song, C. (2020). Spatio-temporal differentia-
476 tion of spring phenology in china driven by temperatures and photoperiod
477 from 1979 to 2018. *Science China Earth Sciences*, 63(10), 1485–1498. doi:
478 <https://doi.org/10.1007/s11430-019-9577-5>
- 479 Yuan, L., Jin, S., & Calabria, A. (2019). Distinct thermospheric mass density
480 variations following the september 2017 geomagnetic storm from grace and
481 swarm. *Journal of Atmospheric and Solar-Terrestrial Physics*, 184, 30-36. doi:
482 <https://doi.org/10.1016/j.jastp.2019.01.007>
- 483 Zhao, T., Liu, P., Zhang, Y., & Ruan, C. (2017). Relating anomaly correlation to
484 lead time: Clustering analysis of cfsv2 forecasts of summer precipitation in
485 china. *Journal of Geophysical Research: Atmospheres*, 122(17), 9094–9106.
- 486 Zirlwagen, D., & Wilpert, K. V. (2010). Upscaling of environmental informa-
487 tion: Support of land-use management decisions by spatio-temporal region-
488 alization approaches. *Environmental Management*, 46(6), 878-893. doi:
489 <https://doi.org/10.1007/s00267-010-9468-4>

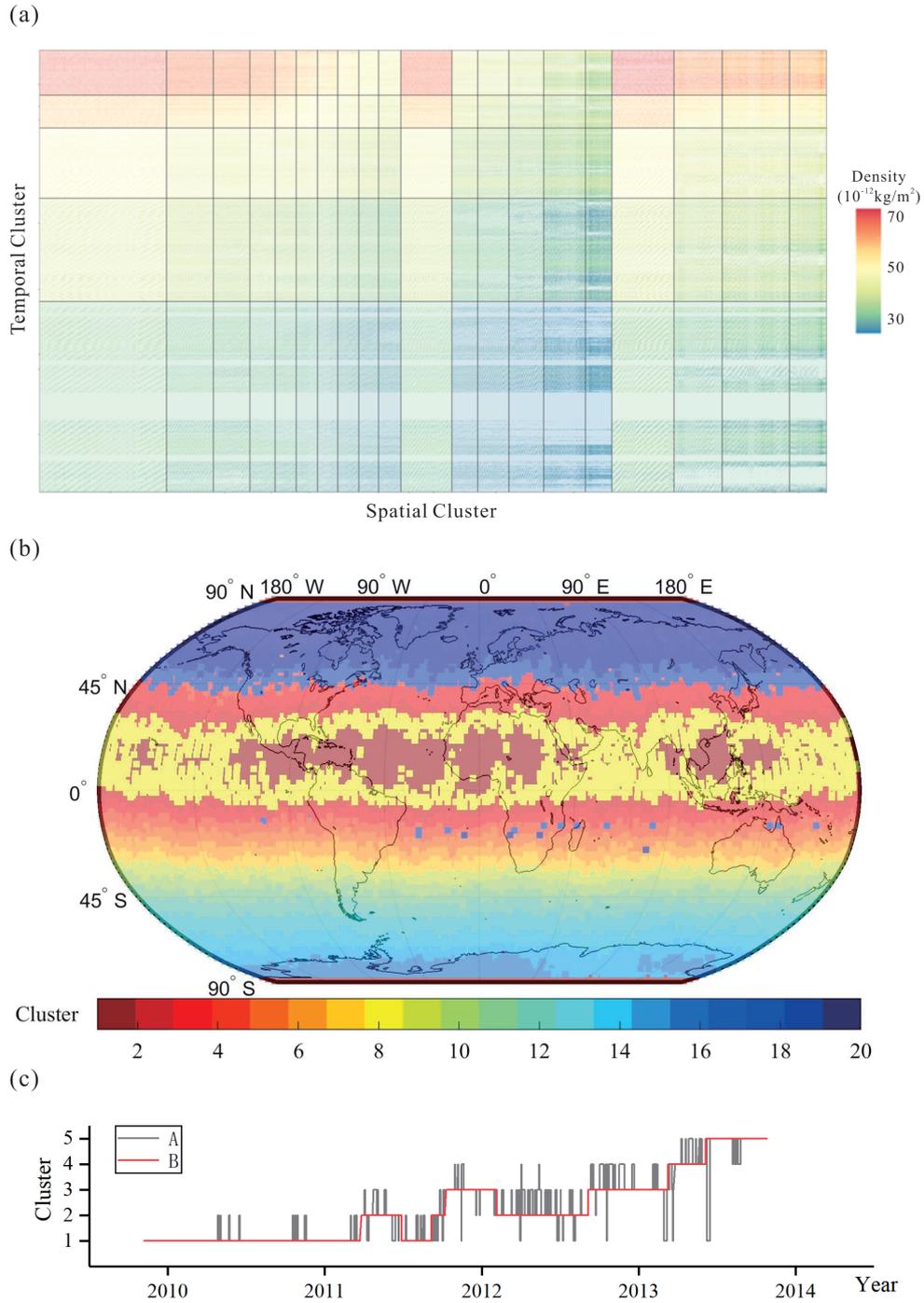


Figure 2. (a) Spatio-temporal co-clustering results of thermospheric mass density, the horizontal axis represents 14940 grids in the dataset, arranged in the order of belonging to the same spatial clusters, the vertical axis represents the 1450 days in the dataset, arranged in the order of belonging to the same temporal clusters, the black solid lines represent the boundary of different groups of time and space, and the color indicates the density value; (b) Spatial clustering result, 1-20 represent 20 groups of spatial clusters; (c) Temporal clustering result, curve A represents original temporal clustering, and curve B represents temporal clustering after removing outliers

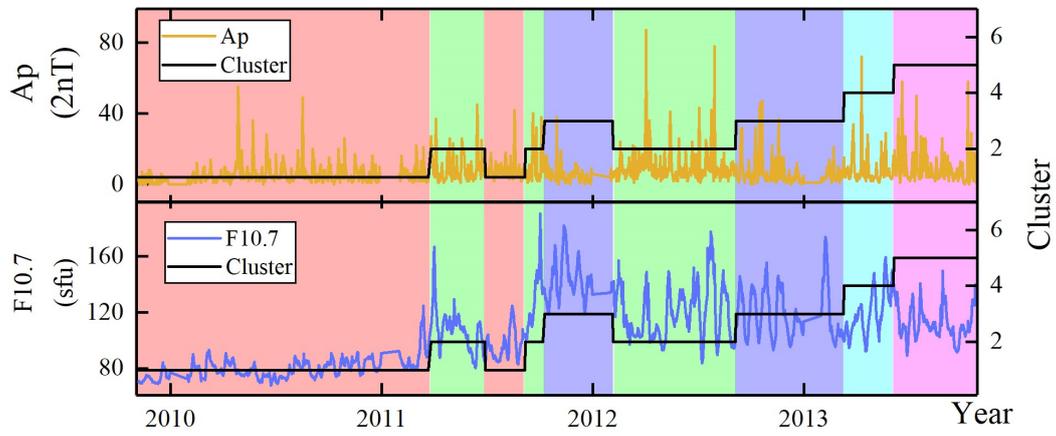


Figure 3. Comparison between temporal patterns and space weather indices